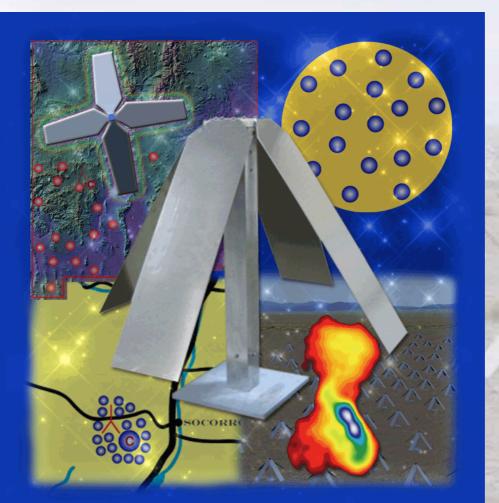


Masaya kuniyoshi (UNM), Sanjay Bhatnagar (NRAO), Greg Taylor (UNM)

(The LWA Project collaboration)



os Alamos

Invent the Future®







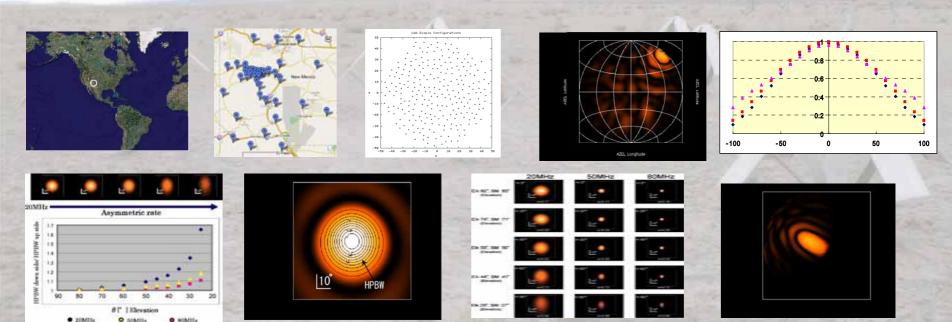
The University of Texas at Austin





Outline

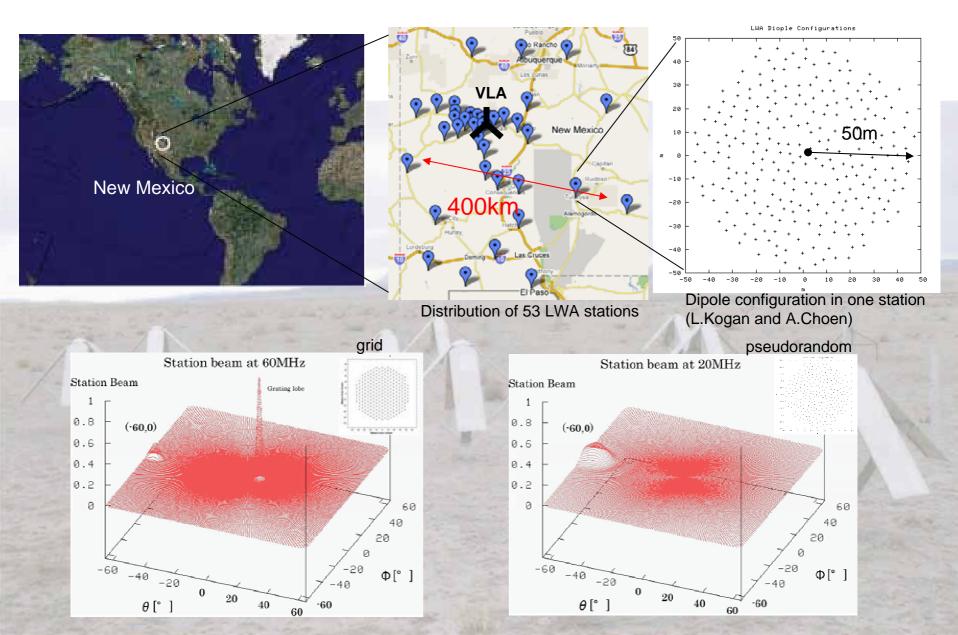
- **1. Long Wavelength Array**
- 2. LWA station beam
 - Elliptical beam
 - Asymmetric beam
 - Pointing error
- 3. LWA imaging simulation
- 4.Summary







Long Wavelength Array (LWA)







Need for a simulator to analyze the LWA

LWA station beam reception pattern

A LWA station beam changes simulator simulator - Function with elevation angles. simulator.close - Function simulator.done - Function

simulator openfromms - Function simulator name - Function simulator summary - Function tor.type - Function imulator settimes - Function imulator observe - Function simulator.setauto - Function simulator.setconfig - Function simulator.setknownconfig - Function simulator setfeed - Function imulator.setfield - Function imulator.setmosaicfield - Function mulator.setspwindow - Function mulator setdata - Function imulator predict - Function mulator setoptions - Function simulator.corrupt - Functio simulator reset - Function imulator.setbandpass - Function imulator setgain - Function imulator setpointingerror - Function simulator setnoise - Function simulator.setpa - Function

simulator setseed - Function

.1 image - Tool nage newimage - Function page newimagefromfile - Function nage imagecale - Function nage.imageconcat - Function nage fromarray - Function page fromascii - Function nage fromfits - Function page fromforeign - Function sage fromimage - Function nage fromshape - Function page.maketestimage - Function nage adddegaxes - Function nage addnoise - Function nage convolve - Function nage boundingbox - Function nage brightnessunit - Function nage.calc - Function nage calcmask - Function nage close - Function nage.continuumsub - Function sage convolve2d - Function nage.coordsys - Function sage coordmeasures - Function nage.decompose - Function sage deconvolvecomponentlist - Function sage remove - Function age.removefile - Functi

1.1 images - Module images

coordsys.conversiontype - Function coordsys.convert - Function coordsys.convertmany - Function measures.observatory - Function coordsys.coordinatetype - Funct measures obslist - Function measures linelist - Function coordsys.copy - Function measures.spectralline - Function coordsys.done - Function measures.sourcelist - Function coordsys.epoch - Function measures.source - Function coordsys.findaxis - Function measures.frequency - Function coordsys.findcoordinate - Functi measures.doppler - Function coordsys.frequencytofrequency measures radialvelocity - Function coordsys.frequencytovelocity - H measures.uvw - Function coordsys.fromrecord - Function measures.tourw - Function coordsys.increment - Function measures expand - Function coordsys.lineartransform - Func measures.earthmagnetic - Functio coordsys.names - Function measures baseline - Function coordsys.naxes - Function coordsys.ncoordinates - Function measures.asbaseline - Function measures listcodes - Function coordsys.observer - Function measures measure - Function coordsys.projection - Function coordsys referencecode - Functi measures.doframe - Function coordsys.referencepixel - Functi measures.framenow - Function coordsys.referencevalue - Funct measures.showframe - Function measures toradialvelocity - Function coordsys.reorder - Function measures.tofrequency - Function coordsys.replace - Function coordsys.restfrequency - Function measures.todoppler - Function coordsys.setconversiontype - Fu measures rise - Function coordsys.getconversiontype - Fu coordsys setdirection - Function measures riseset - Function

CASA

Common Astronomy Software Applications

Many many useful modules

The beam patterns of VLA, ALMA, GBT,

are already in CASA.

measures posangle - Function coordsys.setepoch - Function coordsys.setincrement - Functio measures.separation - Function coordsys setlineartransform - Fu measures addxvalue - Function

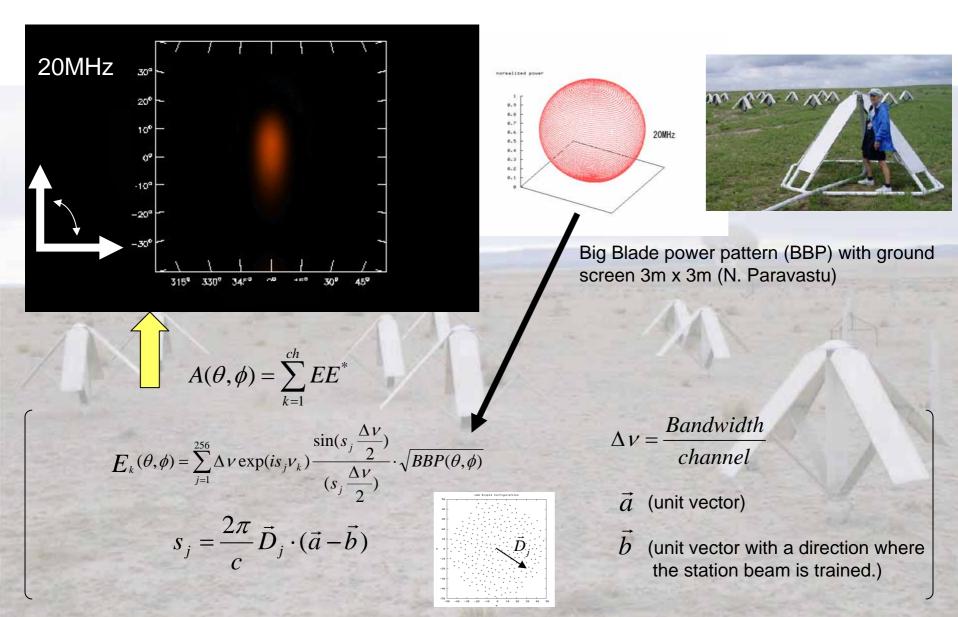
selectrolarization

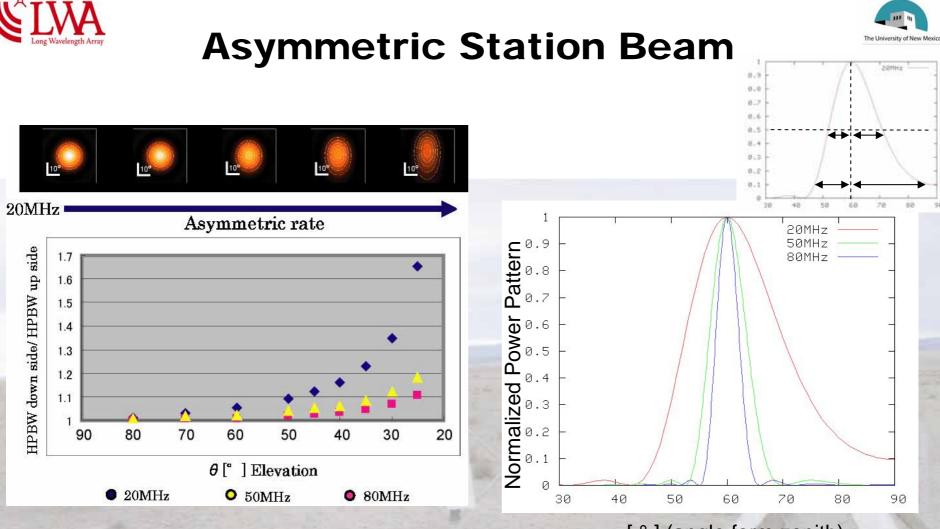
s getdata - Function sputdata - Function s.concatenate - Func s split - Function siterinit - Function siterorigin - Function s.iternext - Function siterend - Function stosdfits - Function s createflaghistory s.saveflags - Function flaglevel - Function s fillbuffer - Function liffbuffer - Function s.getbuffer - Functio s.writebufferflags - Fu measures.torestfrequency - Function sclearbuffer - Function s.continuumsub - Fund suvisf - Function s.ptsrc - Function s.done - Function





Station Primary Beam



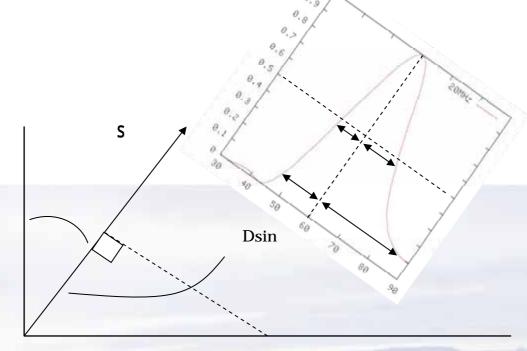


[°] (angle form zenith)

The left graph shows the longitudinal asymmetry of a station beam as a function of elevation at 20MHz, 50MHz and 80MHz. A transverse direction of the beam is always symmetric. A station beam becomes asymmetric as the elevation decreases. The asymmetric effect becomes stronger as the observing frequency becomes lower.







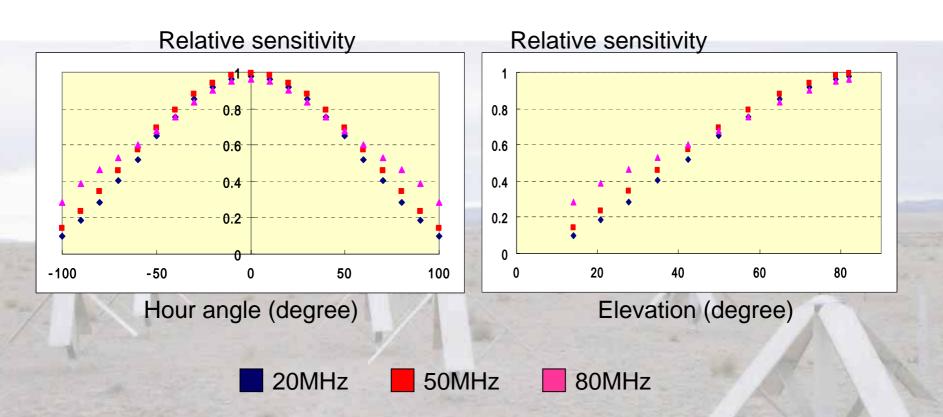
D

As the angle goes from 0 to /2, the value of cos (differentiation of sin) gets smaller. As a result, the beam becomes asymmetric. This effect increases as the frequency decreases.





Sensitivity (example1)

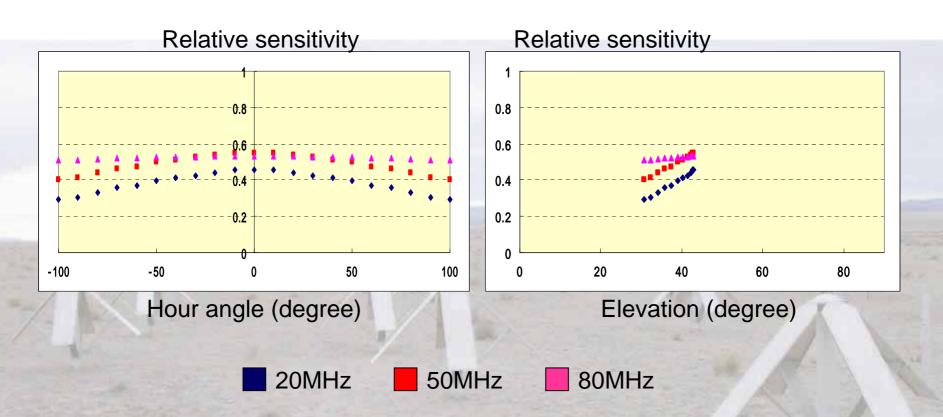


Above shows the simulation results when the LWA Elk station beam (latitude 32.9°) tracks the CygA position (Dec 40.7°).





Sensitivity (example2)



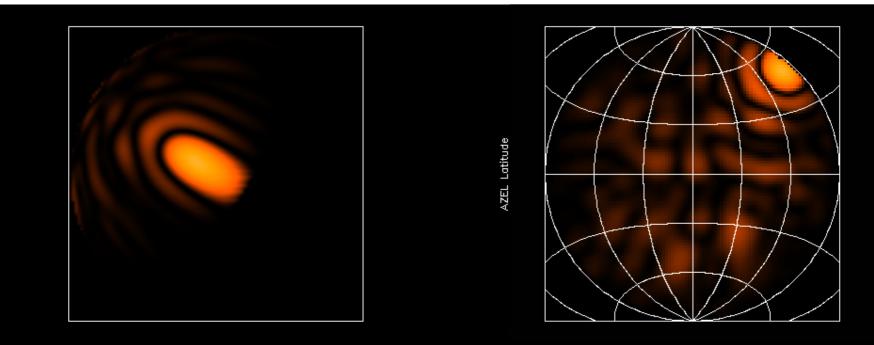
Above shows the simulation results when the LWA Elk station beam (latitude 32.9°) tracks Dec 80° position.





AZEL Latitude

Elk station beam at 20 MHz



AZEL Longitude

Ground coordinate

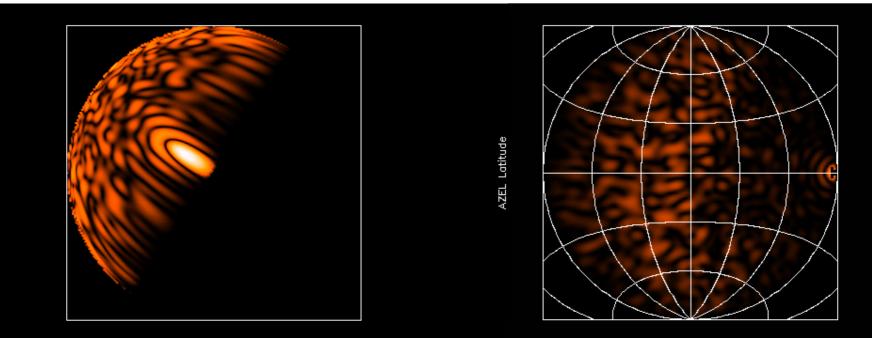
I m coordinate (-100d to +100d)

Above shows the simulation results when the LWA Elk station beam (latitude32.9 $^{\circ}$) tracks the CygA position (Dec 40.7 $^{\circ}$).





Elk station beam at 50 MHz



AZEL Longitude

Ground coordinate

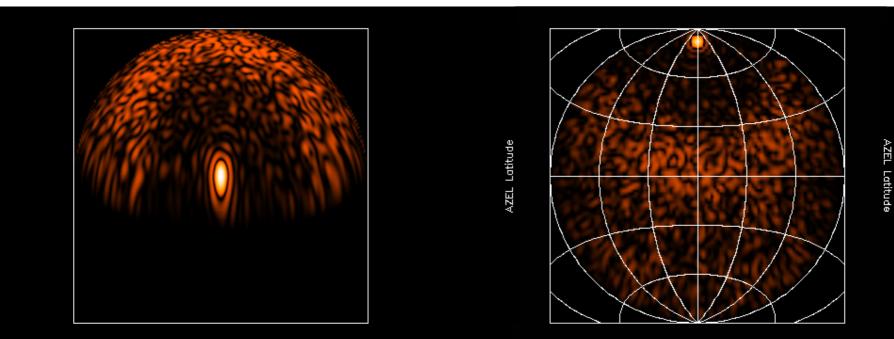
I m coordinate (-90d to +90d)

Above shows the simulation results when the LWA Elk station beam (latitude32.9 $^{\circ}$) tracks Dec 0 $^{\circ}$ position.





Elk station beam at 80 MHz

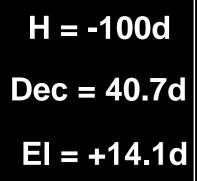


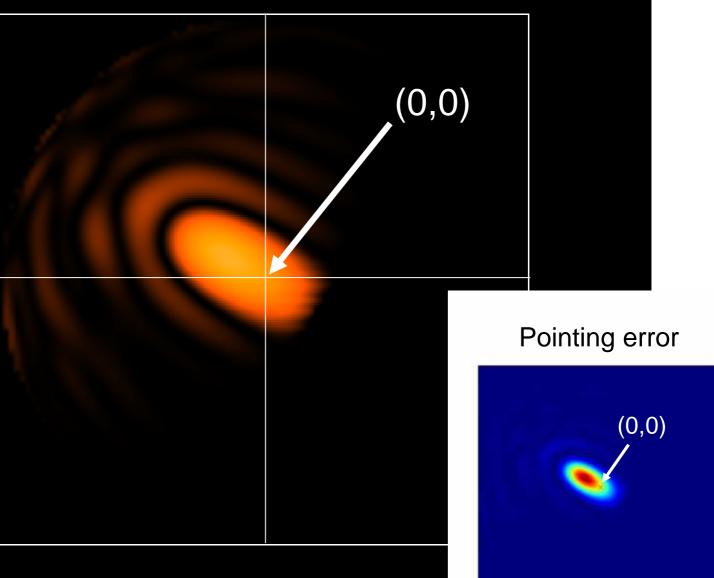
AZEL Longitude

Ground coordinate

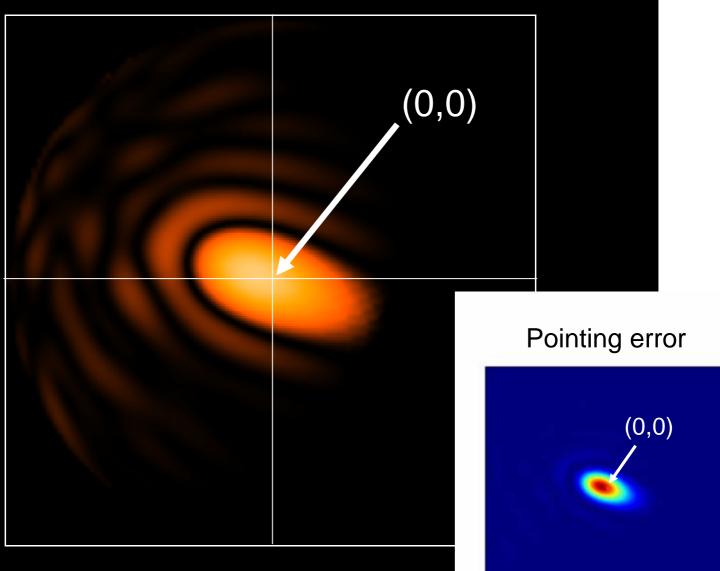
I m coordinate (-180d to +180d)

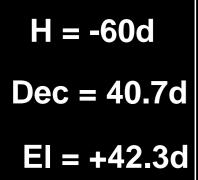
Above shows the simulation results when the LWA Elk station beam (latitude 32.9°) tracks Dec 80° position.

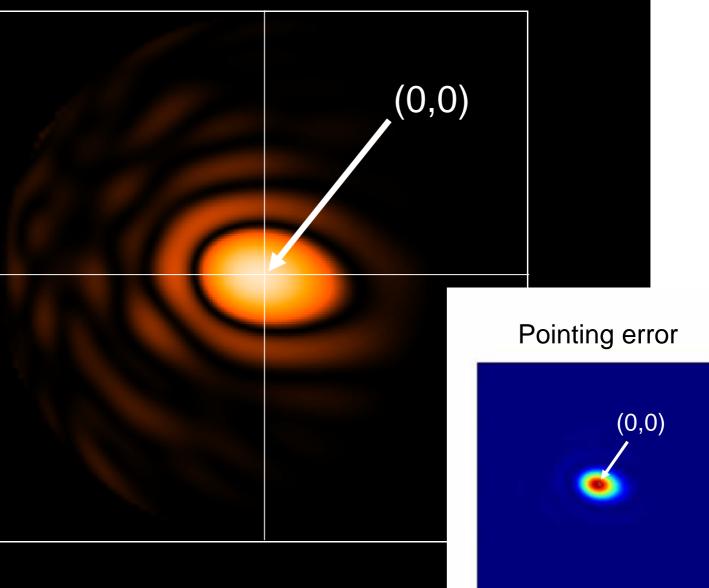


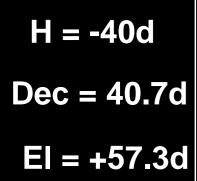


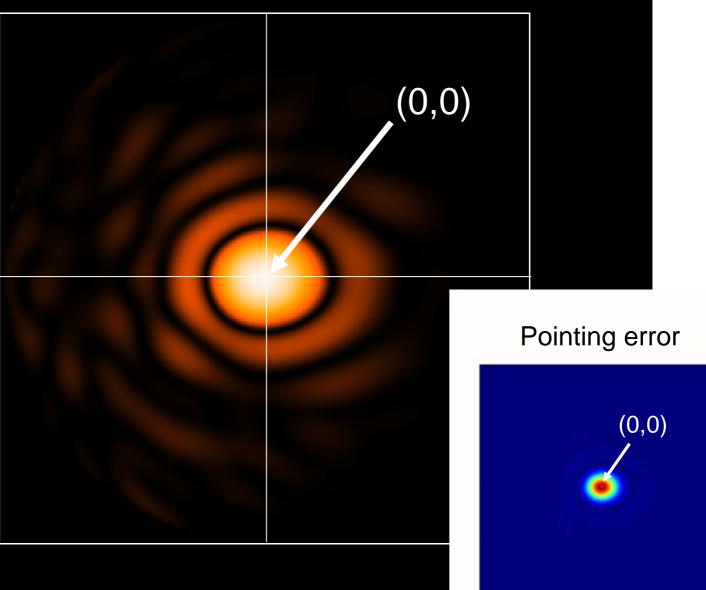
H = -80d Dec = 40.7d El = +27.7d



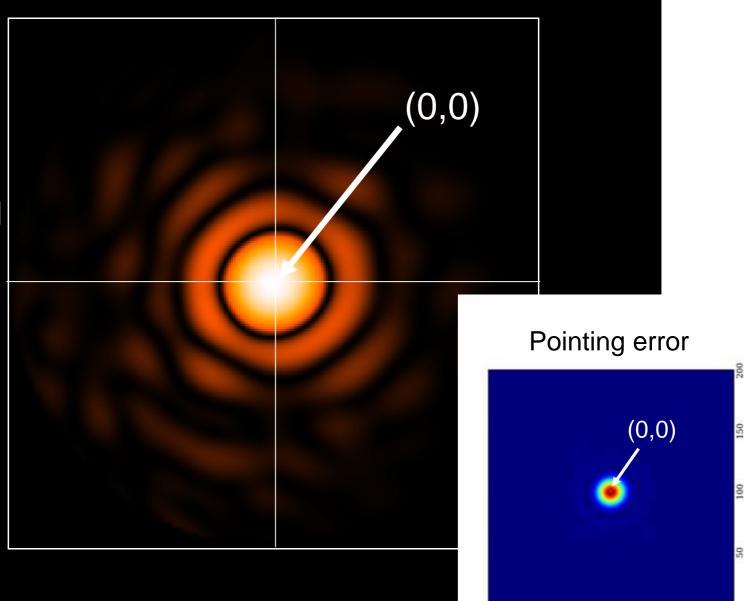




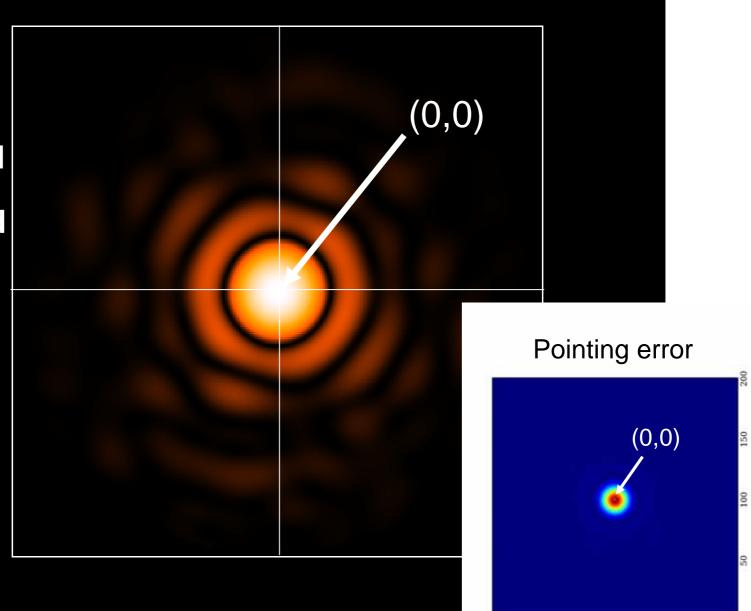




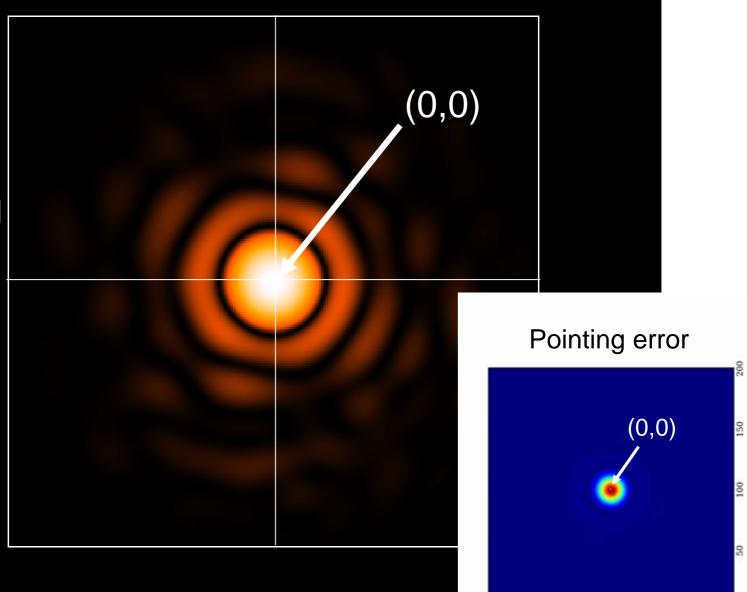
H = -20d Dec = 40.7d El = +72.2d



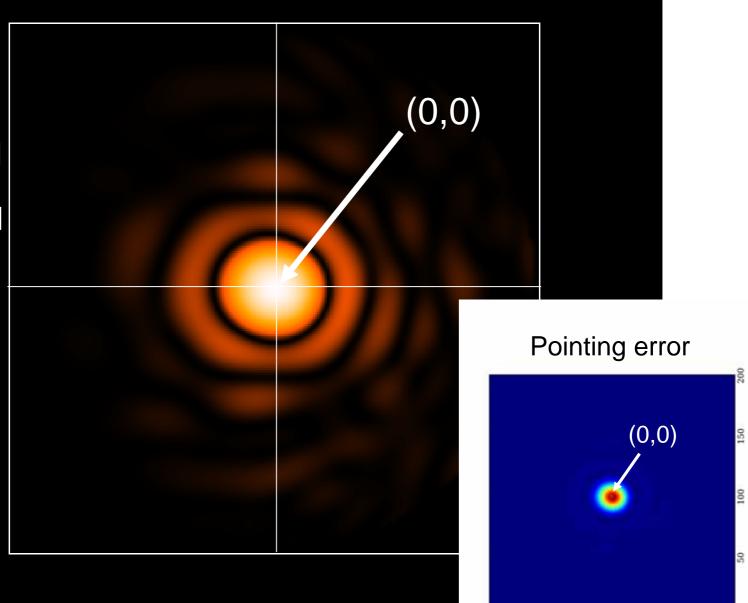
H = 0d Dec = 40.7d El = +82.0d



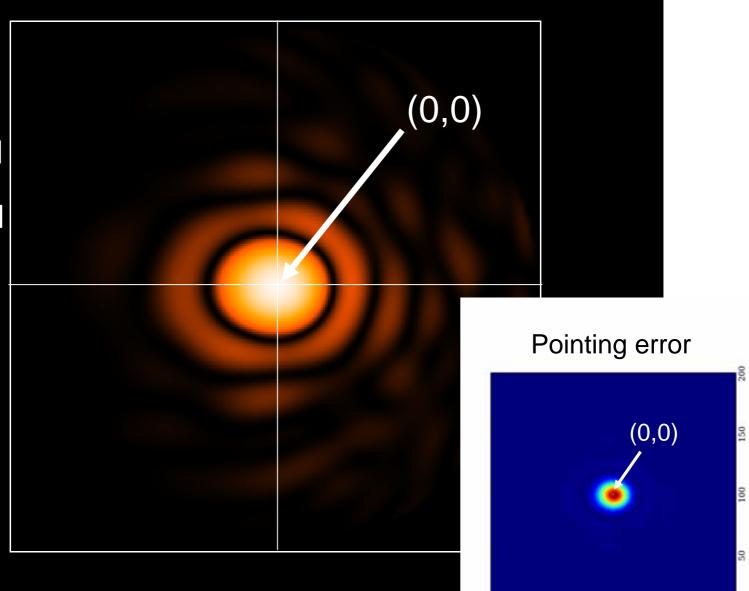
H = +10d $\overline{Dec} = 40.7d$ El = +78.8d

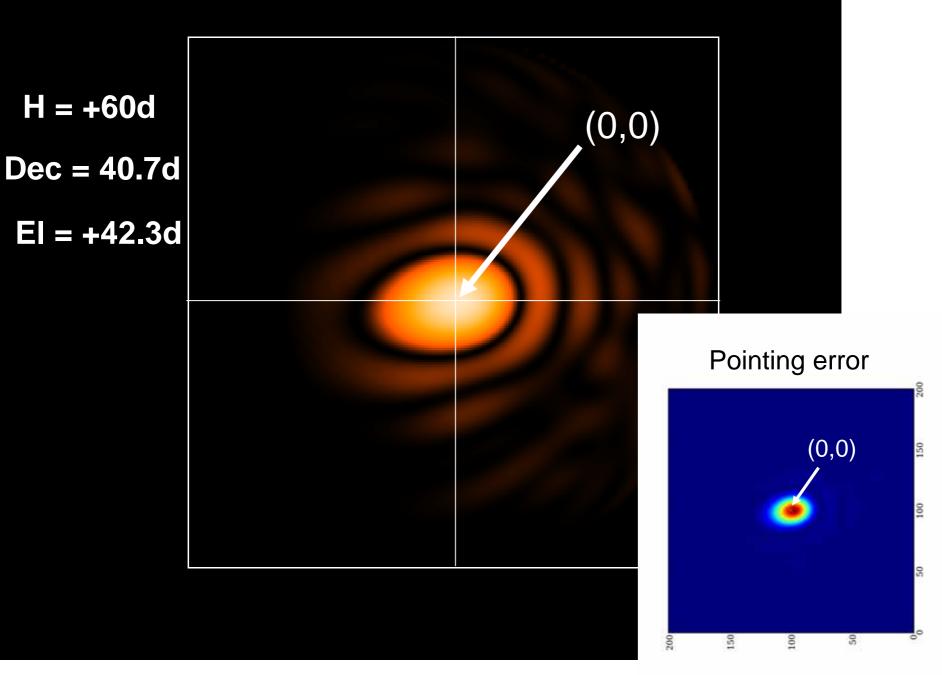


H = +30d Dec = 40.7d El = +64.9d

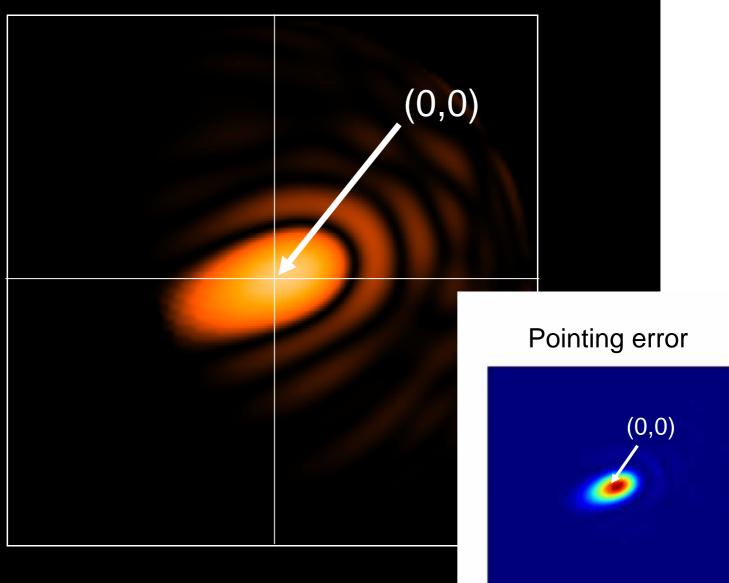


H = +40d Dec = 40.7d El = +57.3d

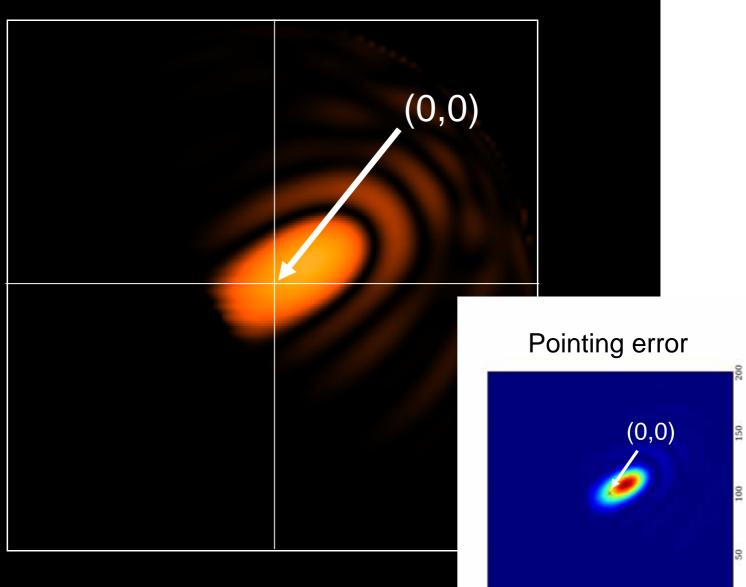




H = +80d Dec = 40.7d El = +27.7d



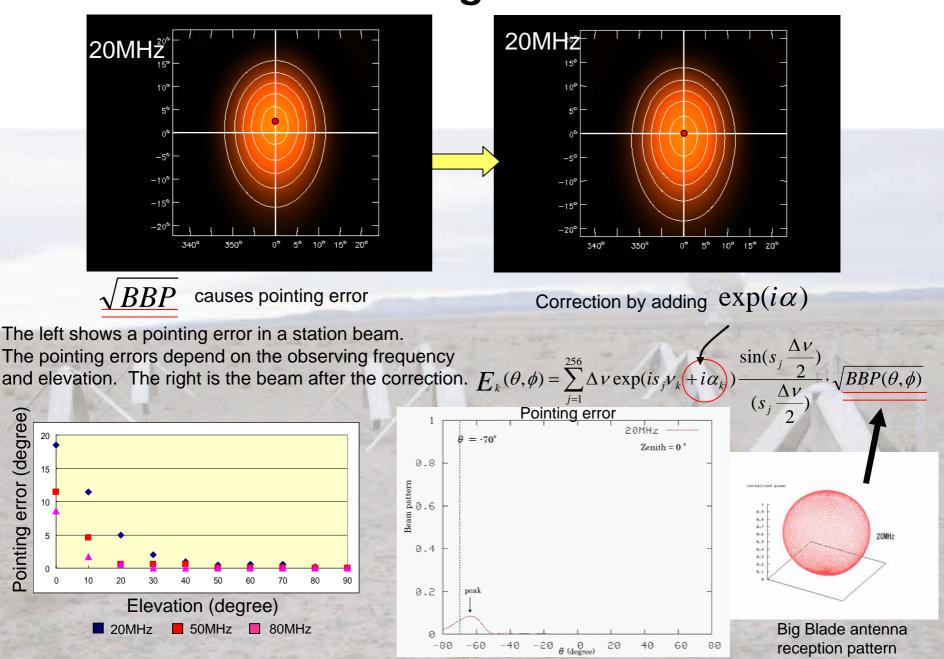
H = +100d Dec = 40.7d EI = +14.1d





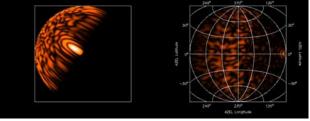
Pointing error

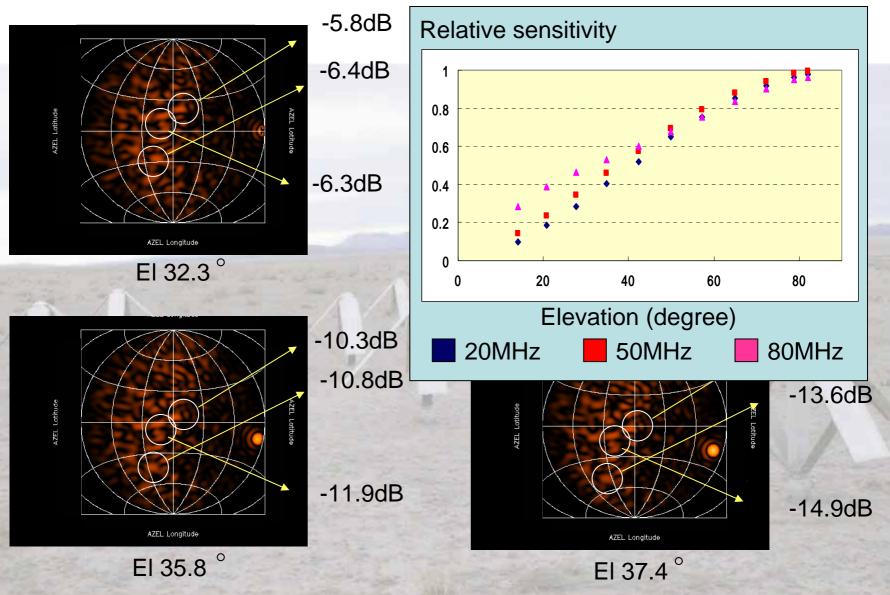




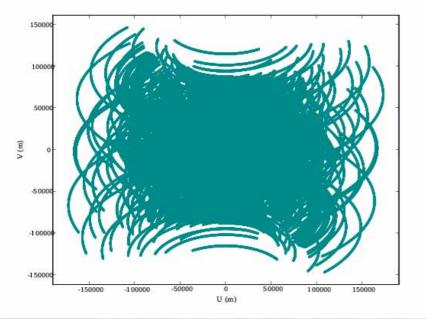


Side lobe at 50MHz





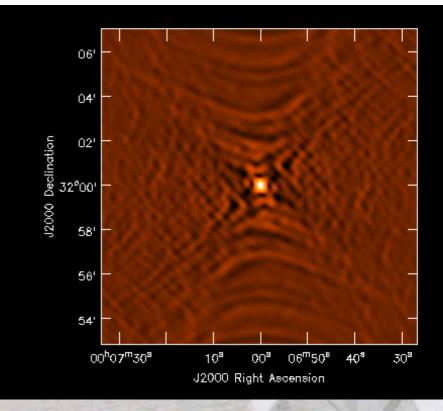




Dec 32.9 degrees [-38°: +38°]



UV coverage and PSF

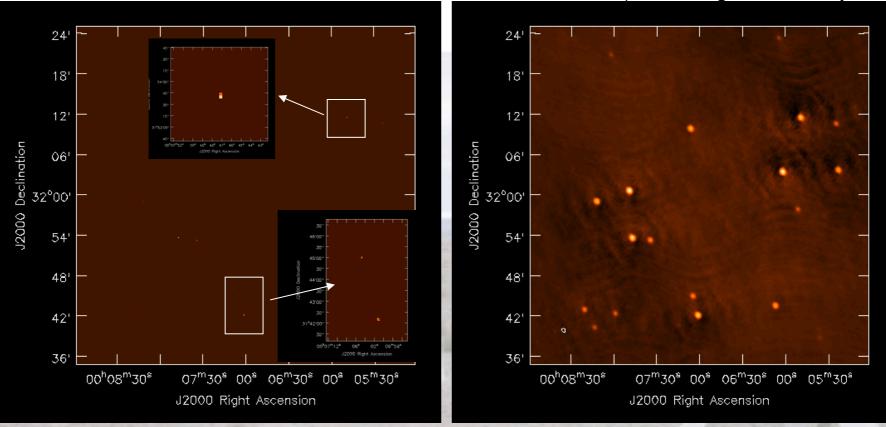






LWA image at 20MHz (preliminary)

(S.Bhatnagar & M.Kuniyoshi)



Simulation model

(Jy/pixel)		
Std Dev	RMS	Mean
5.974e-05	5.974e-05	3.694e-07
Median	Min	Max
0.00	7.868e-05	0.01981

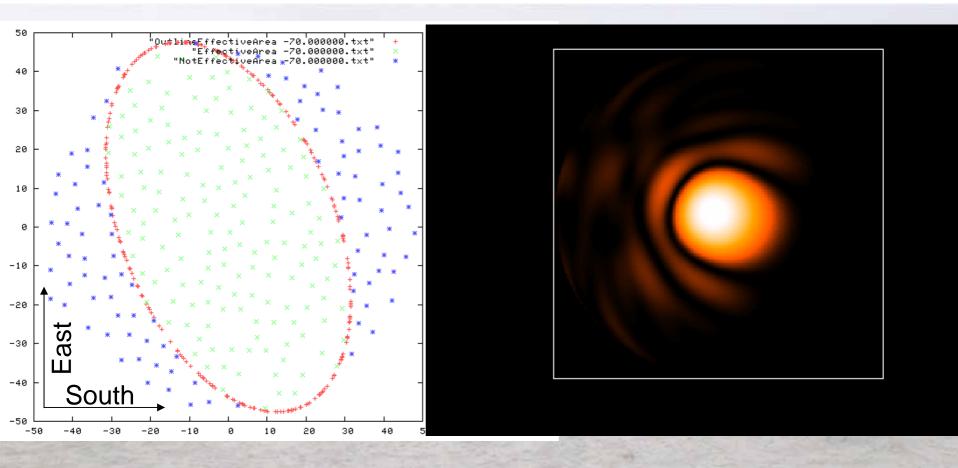
LWA image at 20MHz

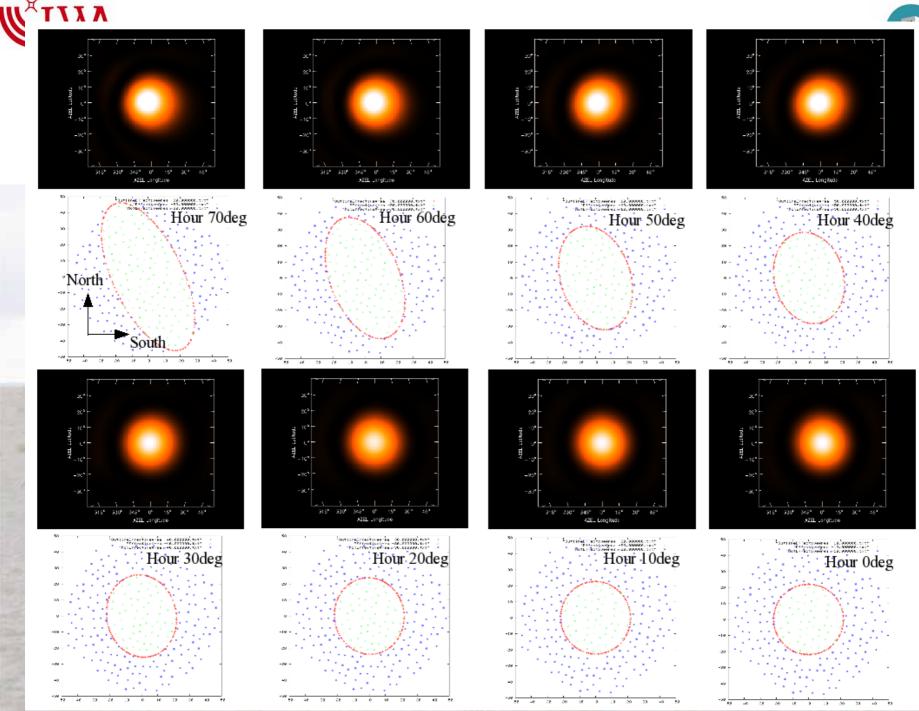
(Jy/beam)		
Std Dev	RMS	Mean
0.0005373	0.0005528	0.0001299
Median	Min	Max
0.0001221	-0.0007181	0.02770



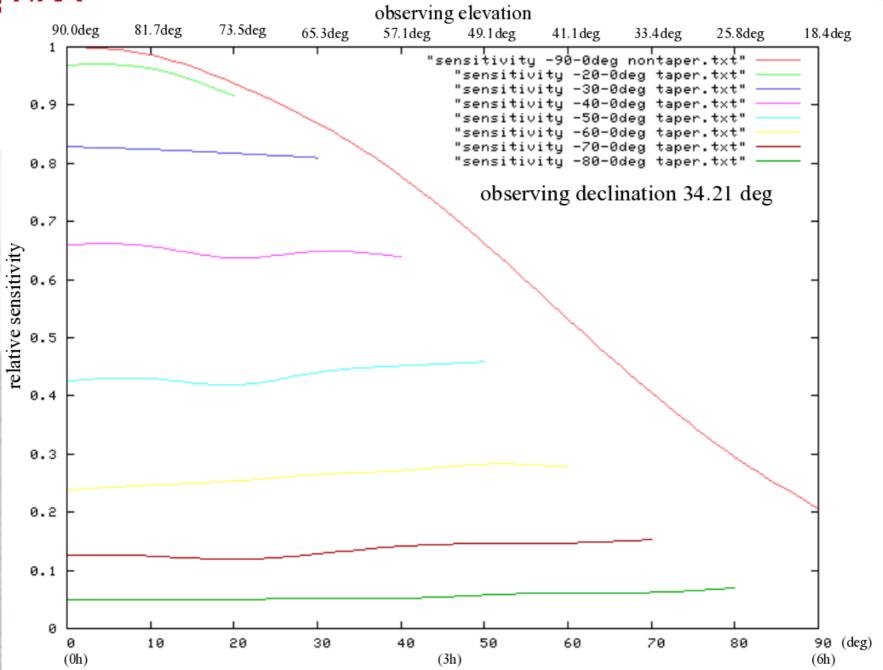


Circular Beam by changing the effective area in a station





AVI I





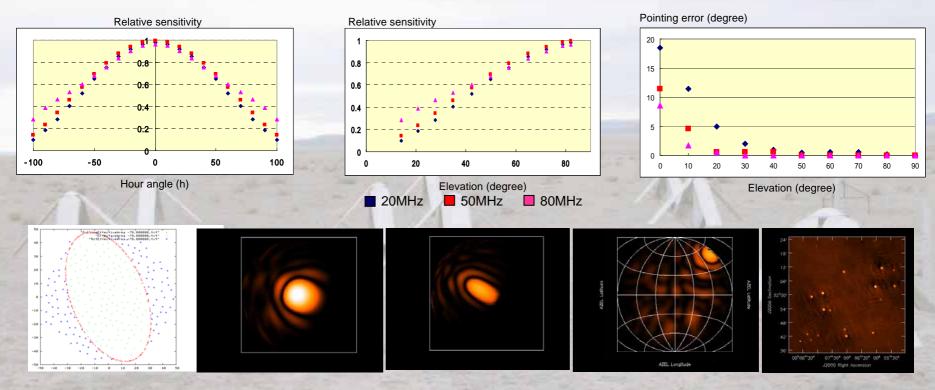


Summary

Station beam

-Pointing error becomes larger with decreasing observing frequency and elevation.

-Sensitivity changes with observing elevation due to the primary beam of the dipole in the station.



Now

We are in the process of creating **more appropriate** images (images from VLSS + 408MHz all sky maps + adding confusion noise) with CASA.





Thank you

